

Gravity waves

Free surface of a liquid in equilibrium in a gravitational field is a plane. If this surface is perturbed then the motion will occur — gravity waves.

○ We shall consider gravity waves in which velocity of the fluid particles is small, such that one can neglect $(\vec{v} \cdot \nabla)\vec{v}$ term compared to $\frac{\partial \vec{v}}{\partial t}$ in Euler's equation. To get an estimate:

If \odot the amplitude of the wave is a , then fluid velocity is $v \sim a/\tau$, where τ is period of the wave. It changes over distances $\sim \lambda$

(wavelength). Thus $(\vec{v} \cdot \nabla)\vec{v} \sim \frac{v^2}{\lambda} \sim \frac{va}{\lambda\tau}$

$$\frac{\partial \vec{v}}{\partial t} \sim \frac{v}{\tau} \Rightarrow (\vec{v} \cdot \nabla)\vec{v} \ll \frac{\partial \vec{v}}{\partial t} \text{ if } \underline{a \ll \lambda}$$

The amplitude should be much smaller than the wavelength.

We also neglect viscosity for the moment ⁽¹⁷⁰⁾

For potential flow of an incompressible liquid $\text{div } \vec{v} = 0$, $\text{rot } \vec{v} = 0 \Rightarrow$

$$\vec{v} = g \text{ grad } \varphi, \quad \nabla^2 \varphi = 0$$

$$\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \vec{v} = -\frac{\nabla P}{\rho} + \vec{g} \Rightarrow$$

$$\vec{v} = 0$$

$$P = -\rho \left(\frac{\partial \varphi}{\partial t} + gz \right)$$

The surface elevation is called $\xi(x, y)$

Then at the surface we have $P = P_0$

$$P_0 = -\rho \left(g \xi + \frac{\partial \varphi}{\partial t} \right)$$

Atmospheric pressure P_0 can be included

by redefining $\varphi \rightarrow \varphi + P_0 t / \rho$

$$\text{Then } g \xi + \frac{\partial \varphi}{\partial t} \Big|_{z=\xi} = 0$$

Vertical velocity at the surface $v_z = \frac{\partial \xi}{\partial t}$ (17)

But for potential flow $v_z = \frac{\partial \varphi}{\partial z}$, thus

$$\left. \frac{\partial \varphi}{\partial z} \right|_{z=\xi} = \frac{\partial \xi}{\partial t} = -\frac{1}{g} \left. \frac{\partial^2 \varphi}{\partial t^2} \right|_{z=\xi}$$

Since the oscillations are small

$z = \xi \Leftrightarrow z = 0$ and we have finally

$$\Delta \varphi = 0$$

$$\left(\frac{\partial \varphi}{\partial z} + \frac{1}{g} \frac{\partial^2 \varphi}{\partial t^2} \right)_{z=0} = 0$$

One needs to supplement these equations

by the boundary condition at the

bottom $\frac{\partial \varphi}{\partial z} = 0$ at $z = -h$.

Deep water (short waves $\lambda \ll h$) ⁽¹⁷²⁾

Consider gravity wave propagating along x .

Let us look for solution in form

$$\varphi = f(z) \cos(kx - \omega t)$$



Then $\nabla^2 \varphi = 0 \Rightarrow$

nam

$$\frac{d^2 f}{dz^2} = k^2 f \Rightarrow f(z) = e^{kz}$$

Boundary condition at the surface

gives us

$$\left(\frac{\partial^2 \varphi}{\partial t^2} + g \frac{\partial \varphi}{\partial z} \right) \Big|_{z=0} = 0 \Rightarrow \omega^2 = gk$$

This is dispersion relation.

For gravity waves it is nonlinear $\omega = g\sqrt{|k|}$

$$\varphi = A e^{kz} \cos(kx - \omega t)$$

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The velocity distribution is obtained from $\vec{v} = \nabla\phi$

$$v_x = -A\kappa e^{\kappa z} \sin(\kappa x - \omega t)$$

$$v_z = A\kappa e^{\kappa z} \cos(\kappa x - \omega t)$$

Velocity diminishes exponentially but rotates

The trajectories of fluid particles are obtained by integrating $\dot{\vec{r}} = \vec{v}$ near

some $\vec{r}_0 = (x_0, z_0)$

$$x = x_0 - \frac{A\kappa}{\omega} e^{\kappa z_0} \cos(\kappa x_0 - \omega t)$$

$$z = z_0 - \frac{A\kappa}{\omega} e^{\kappa z_0} \sin(\kappa x_0 - \omega t)$$

Trajectories - circles



Shallow water (long waves $\lambda \gg h$) (174)

Solving $\Delta \varphi = 0$

$$\left(g \frac{\partial \varphi}{\partial z} + \frac{\partial^2 \varphi}{\partial t^2} \right)_{z=0} = 0$$

with the boundary condition at the bottom $v_z = \frac{\partial \varphi}{\partial z} \Big|_{z=-h} = 0$

we have $\varphi = f(z) \cos(kx - \omega t)$

$$\frac{d^2 f}{dz^2} = k^2 f$$

$f'(z=-h) = 0 \Rightarrow f = \cosh(k(z+h)) \Rightarrow$

$$\varphi(x, z, t) = A \cosh[k(z+h)] \cos(kx - \omega t)$$

Trajectories - ellipses



Boundary condition at the top



$$\omega^2 = k g \tanh(kh)$$



Thus for $kh \ll 1$ linear dispersion



$$\omega = \sqrt{gh} k$$

Viscosity - damping of gravity waves

Consider small viscosity $\lambda^2 \omega \gg \nu$

We still can neglect $(\vec{v} \cdot \nabla) \vec{v}$ term in the Navier-Stokes equation. Then our solution of the potential flow would be good approximation, But it doesn't satisfy boundary conditions at the top (since $\partial_{xz} = -2\eta \phi_{xz}$ and $\partial'_{zz} = -2\eta \phi_{zz} \neq 0$) and at the bottom

$$v_{||} = \frac{\partial \phi}{\partial x} \neq 0 \Rightarrow \text{boundary layers}$$

For small viscosity these layers at the top and at the bottom are thin and can be neglected. (we assume deep water $\kappa h \gg 1$)

The change in energy per unit time is

$$\frac{dE}{dt} = -\frac{\eta}{2} \int \left(\frac{\partial v_i}{\partial x_k} + \frac{\partial v_k}{\partial x_i} \right)^2 dV$$

Introducing potential $\frac{\partial \sigma_i}{\partial x_k} = \frac{\partial^2 \varphi}{\partial x_i \partial x_k} = \frac{\partial \sigma_k}{\partial x_i}$ (176)

$$\begin{aligned} \frac{dE}{dt} &= -2\eta \int \left(\frac{\partial^2 \varphi}{\partial x_i \partial x_k} \right)^2 dV = \\ &= -2\eta \int \left(\varphi_{xx}^2 + \varphi_{zz}^2 + 2\varphi_{xz}^2 \right) dV \end{aligned}$$

The potential $\varphi = \varphi_0 \cos(kx - \omega t + \alpha) e^{kz} \Rightarrow$

averaged over period

$$\overline{\frac{dE}{dt}} = \int_0^{2\pi/\omega} \frac{dE}{dt} dt \frac{\omega}{2\pi} = -8\eta k^4 \int \overline{\varphi^2} dV$$

The average energy is from the virial

theorem

$$\overline{E} = \int \rho \overline{\sigma^2} dV = \rho \int \overline{\left(\frac{\partial \varphi}{\partial x_i} \right)^2} dV = 2\rho k^2 \int \overline{\varphi^2} dV$$

Thus the damping coefficient

$$\gamma = \frac{\dot{\overline{E}}}{2\overline{E}} = 2\nu k^2 = 2\nu \frac{\omega^4}{g^2}$$

Then $E \propto e^{-2\gamma t}$, amplitude of the wave $\propto e^{-\gamma t}$

Dimensional estimates

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Dispersion laws that we derived can be guessed from dimensional reasoning.

Really for deep water the only parameters that enter it are ω , k , and g

$$\circ [\omega] = \frac{1}{\text{sec}} \quad [g] = \frac{\text{m}}{\text{sec}^2} \quad , \quad [k] = \frac{1}{\text{m}} \Rightarrow$$

$$\omega^2 \propto gk$$

Analogously for shallow water one can make [velocity] out of g and h , $c \propto \sqrt{gh}$

$$\circ \omega \propto \sqrt{gh} k$$

Note that the wave speed $c = \sqrt{gh}$ in the relation $\omega = ck$ is equal to the

falling speed from the height h . Also

$$\text{from } \omega^2 \propto gk \quad \Leftrightarrow \quad r = gt^2 \quad \text{thus}$$

circles after the stone ^{has been} dropped in the water expand with the acceleration of the free fall!

Circular waves on the deep water

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We are looking for solutions in the form

$$\psi \propto e^{\kappa z} e^{i\vec{\kappa} \cdot \vec{r} - i\omega t}$$

Then $\nabla^2 \psi = 0$ and from the boundary condition $g \frac{\partial \psi}{\partial z} + \frac{\partial^2 \psi}{\partial t^2} = 0 \Rightarrow \omega^2 = g\kappa$

Initially perturbation was at $r \approx 0 \Rightarrow$

$$\psi(z=0) \propto \int e^{i\vec{\kappa} \cdot \vec{r} - i\omega t} d^2 \kappa \quad \text{with } \omega^2 = g\kappa$$

$$\psi \propto \int e^{i\vec{\kappa} \cdot \vec{r} - i\sqrt{g\kappa} t}$$

To calculate such integrals for large t limit one uses method of

stationary phase

Consider the following integral

$$f(t) = \int_{x_1}^{x_2} dx e^{i t h(x)} \quad \text{in the limit } t \rightarrow \infty$$

1) If $h'(x) \neq 0$ for $x_1 \leq x \leq x_2$ then

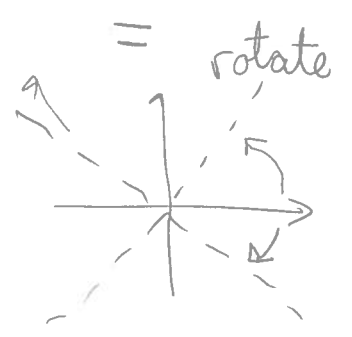
$$\begin{aligned}
 f(t) &= \frac{1}{i t} \int_{x_1}^{x_2} \frac{dx}{h'(x)} \left(e^{i t h(x)} \right)' = \\
 &= \frac{1}{i t} \frac{e^{i t h(x)}}{h'(x)} \Big|_{x_1}^{x_2} + \frac{1}{i t} \int_{x_1}^{x_2} \frac{dx}{h''(x)} e^{i t h(x)} = O(t^{-1})
 \end{aligned}$$

2) If $h'(x_0) = 0$, $h''(x_0) \neq 0$ then

$$f(t) = \int_{x_0-\epsilon}^{x_0+\epsilon} dx e^{i t h(x)} + O(t^{-1})$$

Expanding close to x_0 , $h(x) = h(x_0) + h''(x_0) \frac{(x-x_0)^2}{2} + \dots$

$$\begin{aligned}
 f(t) &= e^{i t h(x_0)} \int_{-\epsilon}^{+\epsilon} ds e^{i t \frac{s^2 h''(x_0)}{2}} \\
 &\approx \frac{e^{i t h(x_0)}}{\sqrt{t |h''(x_0)|}} \int_{-\infty}^{\infty} d\tau e^{i \text{sgn}[h''(x_0)] \frac{\tau^2}{2}} \\
 &= \frac{e^{i t h(x_0)}}{\sqrt{t |h''(x_0)|}} \sqrt{2\pi} e^{i \frac{\pi}{4} \text{sgn}[h''(x_0)]} \propto \frac{1}{\sqrt{t}}
 \end{aligned}$$



Applying it to the integral

$$U \propto \int e^{i\vec{k}\vec{r} - i\sqrt{gk}t} (d^2k) =$$

$$= \int e^{ikr\cos\varphi - i\sqrt{gk}t} k dk d\varphi$$

extremum of the exponent is at

$$\frac{r\cos\varphi}{t} = \frac{\sqrt{g}}{2\sqrt{k}} \Rightarrow k_0 = \frac{gt^2}{4r^2\cos^2\varphi}$$

$$h = \frac{k r \cos\varphi - \sqrt{gk}t}{t} \quad h(k_0)t = \frac{-gt^2}{4r\cos\varphi}$$

$$h''(k_0) = \frac{g^{1/2}}{4k_0^{3/2}} \propto \frac{r^3\cos^3\varphi}{gt^3} \Rightarrow$$

$$U \propto \frac{t^3 g^{3/2}}{r^{7/2}} \int \frac{d\varphi}{(\cos\varphi)^{7/2}} e^{-\frac{igt^2}{4r\cos\varphi}}$$

Performing saddle point for angular

integration ($\cos\varphi \propto 1 - \frac{\varphi^2}{2}$) we obtain

$$U \propto \frac{gt^2}{r^3} \cos \frac{gt^2}{4r} \xrightarrow{n\text{-th}} \text{maximum moves}$$

with $r_n = \frac{gt^2}{8\pi n}$