

**Exercise 1. Canonical quantization of phonons**

We consider a chain of  $N$  atoms of mass  $m$  and coordinates  $x_n$ , with  $n = 1, \dots, N$ . The atoms interact through a potential  $V(x_n) = V(x_1, \dots, x_N)$  that can be written in the *harmonic approximation* as

$$V(u_1, \dots, u_N) = \frac{\lambda}{2} \sum_{n=1}^N (u_{n+1} - u_n)^2 + \frac{1}{2} m \Omega^2 \sum_{n=1}^N (u_n)^2, \quad (1)$$

where  $u_n = x_n - \bar{x}_n$ ,  $|u_n| \ll \bar{x}_n$  measures the (small) deviation from the equilibrium position of each atom  $\bar{x}_n = na$ ,  $a$  being the lattice constant. In (1)  $\lambda$  is the elastic constant of the chain and the  $\Omega$  term constrains each atom at its equilibrium position. The kinetic energy of the atoms is readily written as:

$$T(\dot{u}_1, \dots, \dot{u}_N) = \frac{1}{2} m \sum_{n=1}^N (\dot{u}_n)^2. \quad (2)$$

- a) After writing down the classical Lagrangian of the system and the corresponding equations of motion (Euler-Lagrange equations), solve for the normal modes by imposing periodic boundary conditions (PBC,  $u_n = u_{n+N}$ ). What are the symmetries of the system for  $\Omega = 0$ ? Comment on the resulting spectrum in the two cases  $\Omega \neq 0$  and  $\Omega = 0$ . In the latter case, how does the spectrum look like in the *long-wavelength* regime?

*Hint. Solve the E-L equations with an exponential ansatz  $u_n \propto e^{i(kna - \omega t)}$  and impose PBC to obtain the normal modes  $\omega_l = \omega(k_l)$ . The general solution will look like:*

$$u_n(t) = \sum_{l=1}^N [A_l e^{i(k_l n a - \omega_l t)} + c.c.], \quad (3)$$

where the  $A_l$  are fixed by the initial conditions. The long-wavelength regime is characterized by  $k_l \ll 1$ .

- b) Identify the conjugated momenta  $\pi_n(t)$  and write down the Hamiltonian for the system through a Legendre transformation.
- c) We set  $m = 1$  and introduce the dimensionless operators  $\hat{a}_l = \sqrt{N} \sqrt{\frac{2\omega_l}{\hbar}} \hat{A}_l$ , satisfying  $[\hat{a}_l, \hat{a}_{l'}^\dagger] = \delta_{ll'}$ . Use them to elevate  $u_n(t), \pi_n(t)$  to quantum operators  $\hat{u}_n(t), \hat{\pi}_n(t)$ . Prove the equal-time canonical commutation relation  $[\hat{u}_n(t), \hat{\pi}_{n'}(t)] = i\hbar \delta_{nn'}$ .
- d) Write the Hamiltonian in (c) in terms of the new operators  $\hat{a}_l, \hat{a}_l^\dagger$ . Write a general eigenstate. What is the ground state? What defines an excited state?

**Exercise 2. Planets as blackbodies?**

The Stefan-Boltzmann law states that the emission power per unit surface area of a blackbody reads

$$P_{em} = \sigma T^4 \quad \text{with} \quad \sigma = \frac{\pi^2 k_B^4}{60 \hbar^3 c^2} \approx 5.6704 \cdot 10^{-8} \text{ Js}^{-1} \text{ m}^{-2} \text{ K}^{-4}. \quad (4)$$

- a) Making use of the Stefan-Boltzmann law, estimate the temperature of the Earth, Mars and Venus as if they were blackbodies.

*Hint. The energy emitted and absorbed has to balance.*

- b) The correct results for the average temperatures are 288 K for the Earth, 218 K for Mars and 735 K for Venus. How do they compare with the estimates in (a)? What could be the reasons of the discrepancies?